

## Millimetre-Wave Aperture Synthesis Radiometry for Snow and Ice Mapping

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### INTRODUCTION

Recent years have witnessed a growing interest in passive microwave and millimetre wave imaging for space-based remote sensing of geophysical parameters and features. Applications include measurement and classification of snow and ice cover [1], sea and land surface temperatures, sea surface salinity, soil moisture content and wind velocity [2]. Mechanical, mm-wave pencil beam scanners, such as AMSR and SMMR have previously been employed for earth observation, but for ground resolutions of a few kilometres launching and deploying the large antennas required becomes increasingly problematic. We address here the design of a thinned-array synthetic aperture interferometric radiometer (SAIR) for dual application to (1) mapping of snow and ice cover and (2) determination of ocean wind vector.

The most mature research programmes for Earth observation using radio-frequency aperture synthesis are ESA's MIRAS programme [3] and NASA's ESTAR programme [4] that both operate at L-band. The former uses a two-dimensional Y-shaped array of circular antennas whereas the latter employs a linear array of stick antennas that is push-broom scanned to form a two-dimensional image. Other candidate techniques are the Mills Cross [5] the closely related RadSAR [6] and supersynthesis [7] techniques and hybrid techniques combining a slow conical scan with one- and two-dimensional aperture synthesis such as the so-called QUARTZ technique [8]. Mm-wave SAIR for Earth observation is also pertinent to terrestrial real-time mmw imaging [9,10] for applications such as detection of land mines, foul weather flying aids and surveillance. In seeking to extend these microwave SAIR techniques to mm-wave frequencies it is crucial to consider the ramifications of operation at higher mm-wave frequencies in terms of the increased demands on component specifications and the greatly increased costs.

### COMPARISON OF MM-WAVE SAIR TECHNIQUES

The main challenge in designing a SAIR system is to achieve the technical requirements with acceptable complexity: parameters such as sensitivity and field of view can only be improved at the expense of increased system complexity and cost. It is therefore important that a cost-efficient system is

designed with a performance that is necessary and sufficient for the scientific task. The technical requirements for a SAIR to map snow and ice and to measure ocean wind vector are summarised in Table 1. Table 2 summarises the salient parameters that discriminate between the most pertinent SAIR architectures when designed to meet the snow and ice mapping requirements at 31.5 GHz. Similar differences exist between the systems for operation at 89 GHz. The SAIR architectures parameterised are the Mills Cross, two-dimensional aperture synthesis using a Y array (as for MIRAS) in both a pushbroom and conical scan, one-dimensional aperture synthesis using a conical scan and sub-arrayed antennas (QUARTZ) and one-dimensional aperture synthesis with a pushbroom scan (ESTAR). Dimensions of the antenna arrays vary slightly between the different architectures, but for an angle of incidence of 50° and Hamming weighting with a -6 dB edge taper, the arrays are typically about 3 m parallel to the orbital velocity vector and 1.7 m in the transverse direction.

The use of a conical scan enables the instantaneous field of view to be decreased to reduce electronic complexity at the cost of increased mechanical complexity. The additional use of sub-arraying enables particular small fields of view with efficient suppression of alias responses. The quoted bandwidth  $B$  is either the width of the radio-quiet window or the bandwidth required to achieve the stated sensitivity. The parameters  $N_{\text{correlators}}$ ,  $N_{\text{antennas}}$  and  $N_{\text{receivers}}$  are the required numbers of correlators, antennas and receivers and taken together with the  $B \cdot N_{\text{correlators}}$  product form a measure of the complexity of the SAIR system. It is clear from these data that one-dimensional aperture synthesis with a pushbroom scan offers the highest performance with the lowest complexity. The salient reasons for this is that the minimum number of required pixels are formed and this leads to both reduced complexity and improved sensitivity (radiometric resolution is proportional to the square root of the number of baselines contributing to each pixel [11]). It is also clear that the complexity as defined by the above metrics is an order of magnitude higher for 89 GHz system than for the 31.5 GHz system and one must consider whether this is justified by the increased resolution and the marginal improvement in snow and ice classification.

### SYSTEM DESIGN

In this section we describe the outline design for a polarimetric, mm-wave SAIR. One-dimensional aperture synthesis employs a hybrid technique in which a long linear array of real-aperture

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**Table 1 Technical requirements for ocean wind vector and snow and ice mapping**

Orbit altitude	800 km
Angle of incidence, $\theta$	$\sim 50^\circ$ ( $40^\circ < \theta < 65^\circ$ )
Swath width	>1300 km
Centre operating frequency snow & ice ocean wind vector	31.5 GHz & 89 GHz 31.5 GHz
Available bandwidth	0.4 GHz at 31.5 GHz 2.4 GHz at 89 GHz
Polarisation snow & ice ocean wind vector	Horizontal and vertical Full Stokes parameters
Spatial resolution snow & ice ocean wind vector	6 km at 31.5 GHz 2.5 km at 89 GHz 50 km at 31.5 GHz
Radiometric resolution snow & ice ocean wind vector	1 K 0.05 K
Radiometric accuracy	2 K

stick antennas form over-lapping fan beams on the ground and aperture synthesis within the fan beams enables synthesis of the mm-wave image. Critical aspects for building a high-resolution, mm-wave ESTAR are system calibration and the construction of stick antennas that will be 3 m long and about half a wavelength wide.

#### Mm-wave stick antennas

The most appropriate technology for implementing the stick antennas is a co-phased array of slotted-waveguide leaky-wave antennas. These have less sensitivity to frequency tuning than continuous leaky-wave antennas, lower loss than corporate fed dipole arrays and lower bulk than a co-phased array of pillbox antennas [12].

At 89 GHz the high attenuation of fundamental-mode waveguide limits the length of a simple slotted waveguide

antenna to less than 1 m and frequency tuning of the antenna pattern across the full bandwidth  $B$  would be typically more than five times the width of the antenna pattern of a three metre long antenna. Combining the antennas into a co-phased array reduces the Ohmic losses, whilst the frequency tuning of each element is manifested as aperture plane phase aberration that increases sidelobe levels. To attain sidelobe levels in the real aperture direction that are below -20dB, five elements are required for operation at 31.5 GHz and 20 elements are required at 89 GHz. This is the reason for the increased numbers of receivers shown in columns five and six of Table 2. The aperture efficiency of the antennas is then predicted to be over 84% and broadening of the antenna pattern due to frequency-tuning is less than 0.3% at both frequencies. For polarimetric operation it will be necessary to employ interleaved arrays of horizontally and vertically polarised slotted waveguides [13,14] within the inter-element spacing of 0.82 wavelengths that is necessary to position the first alias responses above the Earth's horizon.

To avoid resonances within the antenna structure the antenna pattern of a slotted waveguide antenna should be displaced from broadside by a few degrees and this requires a small compensatory increase in antenna length to maintain angular resolution. It might be noted that the set of vectors from the array to the far-field then describes a cone rather than a plane. This fact can be exploited to reduce the variation of angles of incidence across the swath width so as to improve the conditioning of the data inversion used for snow and ice classification. As an extreme example, if the stick antennas are co-aligned with the nadir vertical and the beam is tilted to  $43^\circ$  from broadside, the angle of incidence is a constant  $50^\circ$  across the swath width. The penalty however is a 37% increase in length of the antennas, and hence of the number of elements and receivers.

#### Calibration

Calibration has been identified as a major problem for both one-dimensional [15] and two-dimensional [16] Earth-observation SAIRs. It is convenient to consider calibration in two parts: complex transmission from the front-end to the correlators and errors in the transmission function of the antenna array. The former, electronic calibration, can be

**Table 2 Important parameters for the most viable mm-wave SAIR techniques at 31.5 GHz. For ESTAR, the preferred technique, parameters are shown for 31.5 GHz and 89 GHz.**

Parameter	Mills Cross	2D AS		QUARTZ	ESTAR	
		Pushbroom	Conical scan		31.5 GHz	89 GHz
$B$ (MHz)	400	400	400	400	75	445
$N_{\text{correlators}}$	19,602	69,337	121,837	266	832	2044
$B \cdot N_{\text{correlators}}$ (GHz)	7,841	27,735	48,735	233	62.4	910
$N_{\text{antennas}}$	297	322	427	20	25	42
$N_{\text{receivers}}$	297	322	427	100	125	1680
$\Delta T$ (K)	2.4	4.4	2.2	1.5	1	1

accomplished using front-end noise injection [17]. Calibration of the antenna transmission can be accomplished using the G-matrix method [18] which requires the measurement of the impulse response of each antenna pair for the complete field of view - that is to say, it cannot be fully characterised using a single noise source. To achieve a radiometric accuracy of 1 K requires knowledge of the relative physical shape of the antenna array to within one thousandth of wavelength [15,19] and this can be achieved by free-space or optical-fibre interferometry [20,21]. Antenna calibration is thus achieved in two stages: (1) measurement of the full set of antenna pair impulse responses for a measured antenna shape and (2) prediction of the impulse response using modelling and real-time measurement of antenna shape.

#### CONCLUSIONS

An outline design for a dual-band mm-wave polarimetric SAIR has been presented and is considered to be tractable. The envisaged application is for snow and ice mapping and ocean wind vector measurement. The vastly increased complexity of the high-frequency waveband may not be justified by the scientific benefit and operation at an additional lower frequency such as 18 GHz may give improved classification, though with reduced spatial resolution.

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